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**SIMULATIONS OF EMISSION LINES FROM THE NARROW LINE REGION IN SEYFERT GALAXIES**

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General Abstract

One of the biggest questions in astronomy and astrophysics is “How do galaxies form?” Because of the large time scales involved in the formation of galaxies, the only way to learn about the formation of galaxies is through studying galaxies outside the Milky Way by observation and simulation. In the late 20th century astronomers began observing and investigating galaxies that contain supermassive black holes in their center that produce more light than all of the stars within the galaxy, called active galactic nuclei (AGN). When modeling the Narrow Line Region, researchers produce an incident spectral curve based on previous observations represents the spectrum of light produced by the AGN. The SED can be empirically parametrized using spectral indices, which determine the slopes in different areas of the curve. The aim of this research is to synthesize a regression model with data from previous research that will compute all the spectral indices based on one index. We statistically test our regression models using self-written (need better word) Chi Square test code in the Python language. Using the mean values of the spectral indices (Should I include those values?) provided by past research, we craft an initial incident spectral energy distribution curve in the program Cloudy. Next, the spectral indices are varied based on the regression model and the incident spectral curve is “run” through the simulated Narrow Line Region. Based on the shape of the curve, we can infer aspects of the galaxy’s morphology and elemental composition. Research so far has shown that our regression model is statistically significant (taking a leap of faith here) and we have constrained the hydrogen number density, temperature, metallicity, and photon frequency of the incident curve.

Unsure what I should put for our results. At the time of writing this I haven’t been able to go down to the lab, should be able to get there this evening.